



Vicarious Calibration Of The Hyperspectral Imager For Coastal Oceans (HICO) Using MOBY And AERONET-OC Data

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- NOAA Marine Optical Buoy (MOBY)
- Aerosol Robotic NETwork (AERONET) Stations
- US EPA through Blake Schaeffer

Hyperspectral Imager for Coastal Oceans (HICO)



HICO Sensor Parameters

Parameter	Performance	Rationale
Spectral Range	350 to 1070 nm	All water-penetrating wavelengths plus Near Infrared for atmospheric correction
Spectral Channel Width	5.7 nm	Sufficient to resolve spectral features
Number of Spectral Channels	128	Derived from Spectral Range and Spectral Channel Width
Signal-to-Noise Ratio for water-penetrating wavelengths	> 200 to 1 for 5% albedo scene (10 nm spectral binning)	Provides adequate Signal to Noise Ratio after atmospheric removal
Polarization Sensitivity	< 5%	Sensor response to be insensitive to polarization of light from scene
Ground Sample Distance at Nadir	100 meters	Adequate for scale of selected coastal ocean features
Scene Size	50 x 200 km	Large enough to capture the scale of coastal dynamics
Cross-track pointing	+45 to -30 deg	To increase scene access frequency
Scenes per orbit	1 maximum	Data volume and transmission constraints

HICO on Japanese Module Exposed Facility



HICO Processing Activity in APS



Processing Adjustment

- Normalized Water Leaving Radiance (nLw) values derived from:
 - sensor measurement
 - radiometric calibration
 - atmospheric correction algorithm
- Changes occurred in HICO sensor between lab characterizations and installation on ISS
- Sensor calibration degrades over time
- Atmospheric Correction used to derive nLw
- Vicarious Calibration provides updated gains to improve accuracy of data recording / radiometric calibration / atmospheric correction system

Vicarious Calibration Process



Atmospheric Correction Algorithm

- Goal: To retrieve the normalized water-leaving radiance $\binom{n}{k}$ accurately from the spectral measurements of the TOA radiance $L_t(\lambda)$
- Gordon-Wang atmospheric correction algorithm is used in this study
- TOA atmospheric path radiance:

 $L_t = L_{wc} + L_g + L_w + L_r + L_a + L_{ra}$

- Terms represent white-cap, glint, water, rayleigh, aerosol and molecular scattering radiances
- Inverting previous equation to solve for L_w leaves:

 $L_w = L_t - (L_{wc} + L_g + L_r + L_a + L_{ra})$

 Normalized water leaving radiances _nL_w can be computed from L_w and sensor geometry

Radiance Components in Atmospheric Correction

- Lr = f0 * V_{scatter} * pressure
 - f0 = TOA solar irradiance
 - V_{scatter} = Volume Scattering Function
 - Pressure = function of path radiance
 - V_{scatter} and pressure terms depend on sensor/solar geometry
 - Solar irradiance, f0, interpolated to HICO wavelength at HICO bandwidth
- Wavelengths of Lr tables have to match wavelength center and bandwidth of sensor Lt data set



•MODIS-retrieved Lt, Lr, La, and nLw for 412, 443, 488, 531, 547, 667, 748, and 869 (nm) wavelengths at the AERONET-OC location for the Gulf of Mexico, May 4, 2010.

Vicarious Calibration 2 Step Process

Aerosol Scattering Radiance, La

- Emissivity derived from signal response at 748 and 868 nmeter
- Emissivity used to select aerosol model
- Aerosol model used to establish La for processing pixel
- Aerosol model selection process discussed in H.R. Gordon, M. Wang, "Retrieval of water-leaving radiance and aerosol optical thickness over the oceans with SeaWiFS: a preliminary algorithm", Applied Optics January 1994, Vol 33, No 3, Pg 443

Vicarious Calibration requires a 2 step process

- First step generates gains for NIR wavelength bands
- This stabilizes the gains influencing the emissivity derivation
- Second step generates gains for visible wavelength bands

Gain and Offset Computation

- Objective of vicarious calibration is to compute gains and offsets which transform Lt to vLt values that compute insitu nLw values after atmospheric correction is performed
- Gains and offsets can be computed
 - Single date case: ratio of vicarious Lt and measured Lt using notation of "gain = vLt / Lt" where "offset = 0"
 - Multiple date case
 - Use multiple dates to create Lt and vLt pairings
 - Perform linear regression to generate equation (y = mx + b)
 - Let gain = m and offset = b, yields
 - vLt = (gain) Lt + offset
- After gain/offsets are computed scenes are reprocessed using new gain and offset for each band

Current In Situ Data Used for Vicarious Calibration

- Marine Optical Buoy (MOBY) is managed by NOAA
- Moored in uniform water volume near Lanai, Hawaii
- Performs several atmospheric measurements
- Also measures Inherent Optical Properties (IOPs) and Normalized Water-leaving Radiance (nL_w)
- MOBY provides in situ data to perform vicarious calibration for several NASA / NOAA sensors
- MOBY data stored in Ca/Val database for vicarious calibration of hyperspectral data stream





Aerosol Robotic NETwork - Ocean Color (AERONET-OC)

- Managed by NASA Goddard Space Flight Center (GSFC)
- Over 500 locations that record atmospheric data with 14 locations recording in-water data which include:
 - Long Island Sound Coastal Observatory (LISCO)
 - Venice Acqua Alta Oceanographic Tower (AAOT)
- New Gulf of Mexico WaveCIS location managed by NRL
- AERONET data stored in Cal/Val database for vicarious calibration of multispectral HICO_MODIS data stream











Long Island Sound Coastal Observatory (LISCO) Venice, Italy (AAOT)

Representative True Color HICO Scenes



Vicarious Calibration Hyperspectral Verification

- 4 MOBY samples used to train vicarious calibration
- Scatter plot of 4 separate MOBY samples used to test MOBY in situ and HICO nLw values
- Before and after vicarious calibration
- Wavelength locations: 502 and 525 nm





Vicarious Calibration Multispectral Verification

- 8 AAOT samples used to train vicarious calibration
- Scatter plot of 7 separate AAOT samples used to test AAOT in situ and HICO_MODIS nLw values
- Before and after vicarious calibration
- Wavelength locations: 488 and 547 nm





Vicarious Calibration In Situ Data

- Vicarious Calibration process shown for MOBY data has also been performed with AERONET data
 - Gains/offset can be generated for each AERONET station
 - Gains/offsets can be generated for entire set of AERONET stations grouped together
- Insitu data for vicarious calibration can also be provided by collected multi or single date field data

Vicarious Gain/Offset Validation



Pensacola Beach In Situ Data Stations

Vicarious Adjustment: 06/02/11 Pensacola Beach: PB05



No Adjustment: Hyperspectral



Vcal Pensacola Adjustment



Vicarious Adjustment: 06/02/11 Pensacola Beach: PB06



No Adjustment: Hyperspectral



Vcal Pensacola Adjustment



No Adjustment: Multispectral



Vicarious Adjustment: 06/02/11 Pensacola Beach: PB14



No Adjustment: Hyperspectral



Vcal Pensacola Adjustment



No Adjustment: Multispectral



Vicarious Adjustment: 06/02/11 Pensacola Beach: P21



No Adjustment: Hyperspectral



Vcal Pensacola Adjustment



No Adjustment: Multispectral



Future Research Directions

- Improve gain/offset calculation for blue and NIR regions
 - Update solar irradiance to derive Lr more closely with HICO wavelengths
 - Investigate causes for nLw rise in the NIR region
 - Identify aerosol model is selected by vicarious calibration code
- Apply new gain/offset to more scenes and compare with more in situ data
- Apply new gain/offset within automated processing of HICO data

- Performed vicarious calibration for HICO and HICO-MODIS data using training set of MOBY and AERONET data, respectively
- Verified results of updated gains from vicarious calibration using training set by applying them to test set of HICO and HICO-MODIS data
- Validated results of vicariously calibrated gains by matching them with Pensacola Beach in situ data
- Determined additional tasks needed to refine results

Contact Information

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View NRLProcessed HICO data www7331.nrlssc.navy.mil Links

- 1) Browse Imagery
- 2) Mobile Image Viewer App
- 3) HICO Archive Target Search (HATS)

Subscribe for HICO Research Curt Davis cdavis@coas.oregonstate.edu hico.coas.oregonstate.edu



What is HICO?

Log In Become a HICO Data Use

Data

age Gallerie:

rrent Project

The Hyperspectral images for the Costati Ocean (HICC¹⁰) is an imaging spectrometer based on the PHILS alloher imaging spectrometers. HICO is the first spaceborn imaging spectrometer designed coverage (380 to 590 nm sampled at 57 nm) and a very high signat-broke stato to resolve the complexity of the costati ocean. HICO is showed by the <u>Office of Assul Resource</u> as an innovative Naval Prototype (INP), and will demonstrate costati products including water clarity, bottom types. Latitymethy and combine vigetation maps. As an INP, HICO allo sponsore is a clarable to empetity of the space mission by abgring provine PHILS introduces including water clarity, acrossitive ways to reduce the cost and schedule of this space mission by abgring provine PHILS across introduces models by the <u>office</u> possible acrossitive provides the cost and schedule of this space mission by abgring provine PHILS acrossitive schedule of the space mission by abgring provine PHILS acrossitive the cost and schedule of the space mission by abgring provine PHILS acrossitive the possible acrossitive the possible acrossitive the cost and schedule of the space mission by abgring provine PHILS acrossitive the cost and schedule of the space mission by abgring provine PHILS acrossitive the possible acrossitive th

HICO Status



The HICO program was initiated in February 2005. In January 2007 HICO was selected to fly on the Japanese Experiment Modele Expected Facility (JEN-EP) on the International Space Station. Construction began following the Critical Design Review on November 15, 2007. HICO was completed in July 2008 and it was becompeted in High 2008 and it was becompeted in High 2008 and it was becompeted in July 2008 and it was also neeved support from NASA and on the JEN-EF. HREP was isunched on the 1-2 Transfer Vehicle (HYT) Sectember 10.

Questions

Thank you for your interest in this project

